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On the Mutual Relations Between the Orbits of Certain Asteroids.

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(Read before the American Philosophical Society, September 2, 1S92.)
The present writer, several years since,* called attention to the fact that in some parts of the asteroid zone the orbits of particular members have a striking resemblance to each other. These significant coincidences have been regarded by astronomers as worthy of study, and, in addition to the binary and ternary clusters pointed out by the writer, others have been designated by Tisserand, $\dagger$ of Paris, and by Monck, of Dublin. $\ddagger$ These groups, according to the former, cannot be regarded as chance arrangements. "A glance at the list," says Mr. Monck, "will show that the resemblance frequently extends beyond a single pair and embraces what may be called a family-a circumstance which is known to occur in the case of comets also." The writer's list (which might be extended) is as follows :

## Grours of Asteroids.

|  |  | Names. | $a$ | e |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| I |  | Huberta | 3.4580 | 0.1103 | 60 | $10^{\prime}$ | 3290 | 45 |
|  |  | Hermione | 3.4535 | 0.1255 | 7 | 36 | 357 | 36 |
| II. | (106) | Dione. | 3.1670 | 0.1788 | 4 | 38 | 25 | 57 |
|  | (104) | Clymene | 3.1560 | 0.1407 | 2 | 53 | 62 | 30 |
|  | (171) | Ophelia | 3.1554 | 0.1142 | 2 | 33 | 148 | 31 |
|  | (62) | Erato | 3.1241 | 0.1756 | 2 | 12 | 39 | 0 |
|  | (287) | Silesia. | 3.1190 | 0.1217 | 3 | 40 | 65 | 16 |
|  | (212) | Medea | 3.1157 | 0.1013 | 4 | 16 | 56 | 18 |
|  | (86) | Semele. | 3.1015 | 0.2193 | 4 | 47 | 20 | 10 |
|  | (305) |  | 3.0973 | 0.1927 | 4 | 26 | 104 | 37 |
|  | (34.5) | Vera | 3.0966 | 0.1975 | 5 | 11 | 27 | 48 |
|  | (223) | Rosa. | 3.0937 | 0.1206 | 1 | 59 | 108 | 35 |
|  | (268) | Adurea. | 3.0853 | 0.1285 | 2 | 25 | 184 | 48 |

[^0]$\dagger$ Annuaire, 1891.
$\ddagger$ Sid. Mess., October, 1888, p. 334.
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| Names. |  |  | $a$ | $e$ | $i$ |  | $\pi$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $\{$ (238) | Hypatia. | 2.9081 | 0.0876 | 120 | $23^{\prime}$ | $28^{\circ}$ | $24^{\prime}$ |
|  | (191) | Kolga...... | 2.8967 | 0.0876 | 11 | 20 | 23 | 21 |
|  | (1) | Ceres | 2.7673 | 0.0763 | 10 | 37 | 149 | 38 |
|  | (237) | Cælestina | 2.7607 | 0.0738 | 9 | 46 | 282 | 49 |
|  | (116) | Sirona...... | 2.7669 | 0.1433 | 3 | 35 | 152 | 47 |
|  | $\{$ (278) | Paulina.... | 2.7575 | 0.1331 | 7 | 50 | 199 | 52 |
|  | (213) | Lilaea. | 2.7563 | 0.1437 | 6 | 47 | 281 | 4 |
| VI. | (206) | Hersilia.... | 2.7390 | 0.0389 | 3 | 46 | 95 | 44 |
|  | (203) | Pompei .... | 2.7376 | 0.0587 | 3 | 13 | 42 | 51 |
|  | (160) | Una........ | 2.7287 | 0.0624 | 3 | 51 | 55 | 57 |
|  | (301) | Bavaria | 2.7258 | 0.0660 | 4 | 53 | 24 | 4 |
| VII. | (97) | Clotho | 2.6708 | 0.2550 | 11 | 46 | 65 | 32 |
|  | (3) | Juno | 2.6683 | 0.2579 | 13 | 1 | 54 | 50 |
| VIII. | (249) | Asporina... | 2.6947 | 0.1050 | 15 | 38 | 256 | 6 |
|  | (218) | Bianca.... | 2.6653 | 0.1155 | 15 | 13 | 230 | 14 |
| 1 N. | $\{(66)$ | Maia | 2.6454 | 0.1758 | 3 | 6 | 48 | 8 |
|  | (37) | Fides | 2.6440 | 0.1750 | 3 | 7 | 66 | 26 |
| X. | \{ (193) | Ambrosia | 2.5758 | 0.2854 | 11 | 38 | 70 | 52 |
|  | (134) | Sophrosync | 2.5647 | 0.1165 | 11 | 36 | 67 | 83 |
|  | \{ (i9) | Eurynome . | 2.4436 | 0.1945 | 4 | 37 | 44 | 22 |
|  | (19) | Fortuna | 2.4415 | 0.1594 | 1 | 33 | 31 | 3 |
| XII. | (249) | Ilse | 2.3793 | 0.2195 | 9 | 22 | 14 | 16 |
|  | $\{(115)$ | Tliyra. | 2.3791 | 0.1939 | 11 | 35 | 43 | 2 |
|  | (84) | Clio | 2.3629 | 0.2360 | 9 | 40 | 330 | 20 |

Noten-a, $e, t$ and $\bar{a}$ represent the distances, eceentricitles, inclinations and peribelia respectively.

## Remarks.

1. The second eluster has eleven known members, the average inclina. tion being about $3035^{\prime}$, that is, no one differs from the mean ns much as 20. Of the other groups, severnl are not less striking in the closeness of their relutlons.
2. When the earth, as well as Mars itself, was yet a part of the solar ntmosphere, these individual phanetolds were starting on the puths prescribed them. Into how many nebulous fragments they may have been subdivided, and to what extent these ramiflentions miny yet be traced, let the astronomer of the future inquite.

[^0]:    *1887.

